

Solving Mysteries Using Cloudy

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Some Background







AGN § T O R M Project (2014)















Reverberation Mapping



Estimation of AGN's Mass

$$M = \frac{rv^2}{G}$$

Two parameters needed:

1-**velocity** \rightarrow from Doppler line broadening!

2-**distance** → through "Reverberation Mapping"



NGC 5548

Distance: 245 million LY

Mass of BH: $7 \times 10^7 M_{\odot}$





What they expected to see:



Also:

The dimensionful ionization parameter ξ is defined as:

$$\xi = \frac{L}{n_H R^2} [\text{erg cm s}^{-1}]$$

• The dimensionless ionization parameter U is defined as:

$$\mathbf{U} = \frac{Q_H}{4\pi c n_H R^2}$$

Also:



Part I:

The Sad Story 2014-2017





Results from the campaign

1- Emission Line Holiday (STORM)



180 days of observations-2014

The Continuum **CIV Emission Line**

Goad et la. 2016









3- Heavy Absorption in X-ray (Anatomy)



Mehdipur et al. 2015



Why are these results so important?



Continuous spectrum



The correlation between the continuum and the emission lines is the basic requirement for the RM method







Mass of BH is what everybody wants to know. It controls the galaxy and it teaches us about the evolution of the galaxies.

Part II:

Cloudy Team Into Action









Dehghanian et al. 2019a





Cloudy:

1- Takes the SED produced by the AGN

2- Passes this SED through the obscurer and / predicts the transmitted SED

3-Passes the OBSCURED SED through the absorption component1 and predicts the spectrum observed by HST





Part III:

The Happy Ending

2019-2020





The Absorption-Line Holiday





There are two possibilities

1- Holiday is a result of variable luminosity

2-Holiday is a result of variable shape



Table SED "NGC5548.sed" set save prefix "obs1" hden 10 xi -1.2 stop column density 22.08 save continuum units kev ".con" save transmitted contnuum ".tran"



```
Table Read "obs1.tran"
nuF(nu) 3.023 0.2
set save prefix "cf99"
hden 4.72
#stop zone 1
stop column density 21.5
save line list ".lin" "lines.dat" absolute no hash
save continuum units kev ".cone"
save continuum units angstorm ".cona"
save species column densities ".dens" no hash last
"H"
"H+"
"H2+"
"C"
"C+"
"C+2"
"C+3"
"Si"
"Si+"
"Si+2"
"Si+3"
"N"
"N+"
"N+2"
"N+3"
"N+4"
"He"
"He+"
"He+2"
"He[2]"
end
```



There are two possibilities

1- Holiday is a result of variable luminosity

2- Holiday is a result of variable shape







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"Si+3"
"N"
"N+"
"N+2"
"N+3"
"N+4"
"He"
"He+"
"He+2"
"He[2]"
end
```



Changes in the obscurer affects the shape of the SED



Line of Sight Covering Factor 100 %



Table SED "NGC5548.sed" set save prefix "obs1" hden 10 xi -1.2 stop column density 22.08 save continuum units kev ".con" save transmitted contnuum ".tran"







The effects of variable covering factor on the shape of the SED



Dehghanian et al. 2019a



The effects of variable CF on the absorption lines



Dehghanian et al. 2019a







The Emission-Line Holiday





The Emission-Line Holiday



Dehghanian et al. 2019b

There are two possibilities



2-Holiday is a result of variable wind

Changes in the BLR- considering the equatorial obscurer



Dehghanian et al. 2019b









The Emission-Line Holiday



Part IV:

Lets Go Beyond







Table SED "NGC5548.sed" set save prefix "LOC_case1" hden 13 vary grid 10 18 0.25 phi(H) 22 vary grid 20 24 0.25









A Novel Approach to Trace the Disk Wind



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A Novel Approach to Trace the Disk Wind



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Temperature [K]



C IV (EW_{wind}=EW_{BLR}= 50% EW_{total})



Fe K α (EW_{wind}=EW_{BLR}= 50% EW_{total})







Atlas of UV and X-Ray Spectroscopic Signatures of the Disk Wind



Atlas of UV and X-Ray Spectroscopic Signatures of the Disk Wind



Dehghanian et al. 2021

Future:

Not a unique phenomenon



AGN STORM2: Mrk 817



Thank You for Your Attention!

Any questions? Please ask.

